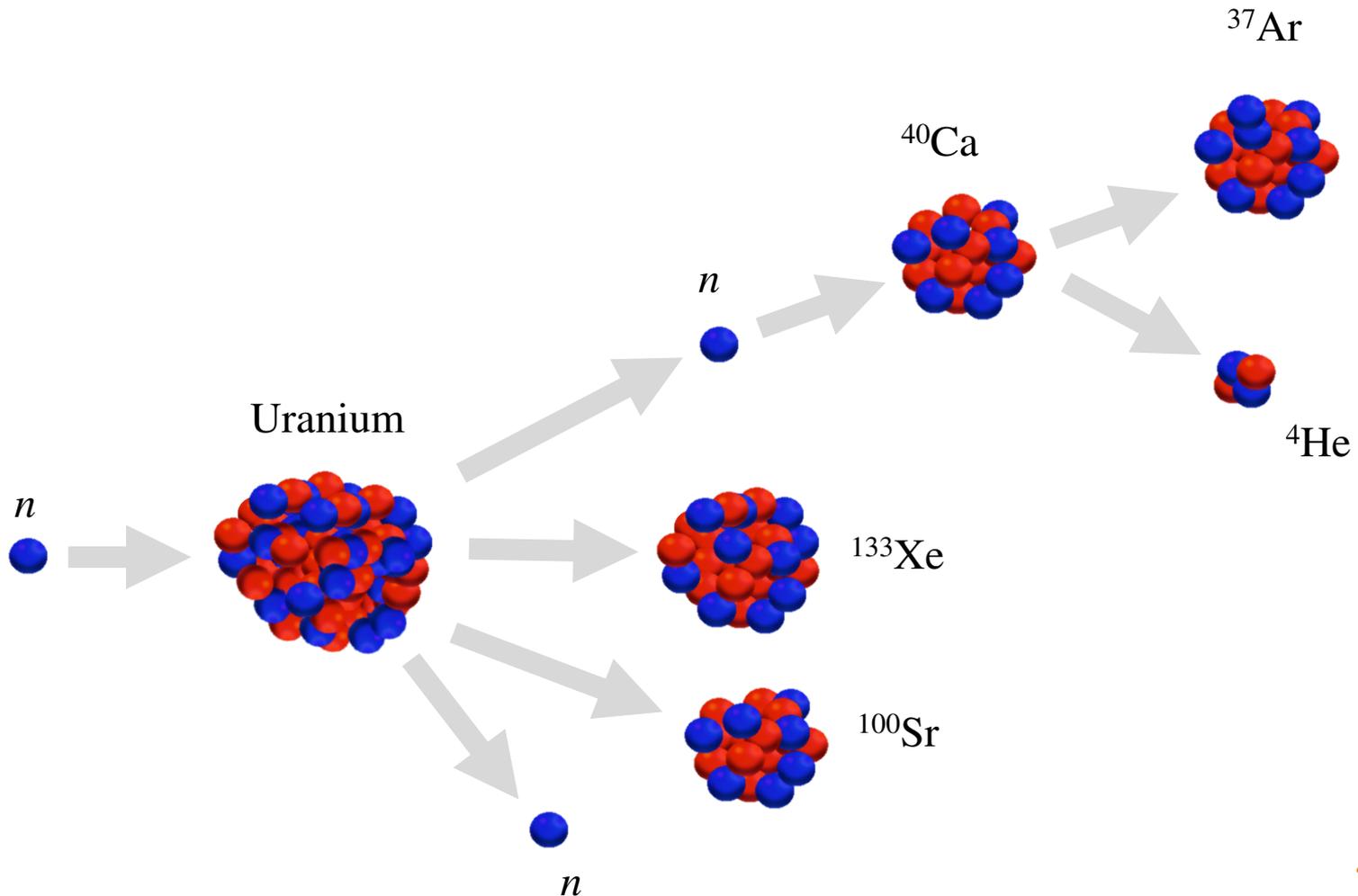


The use of Argon-37 measurements as a coincident signature with radioxenon and in high background environments in the International Monitoring System

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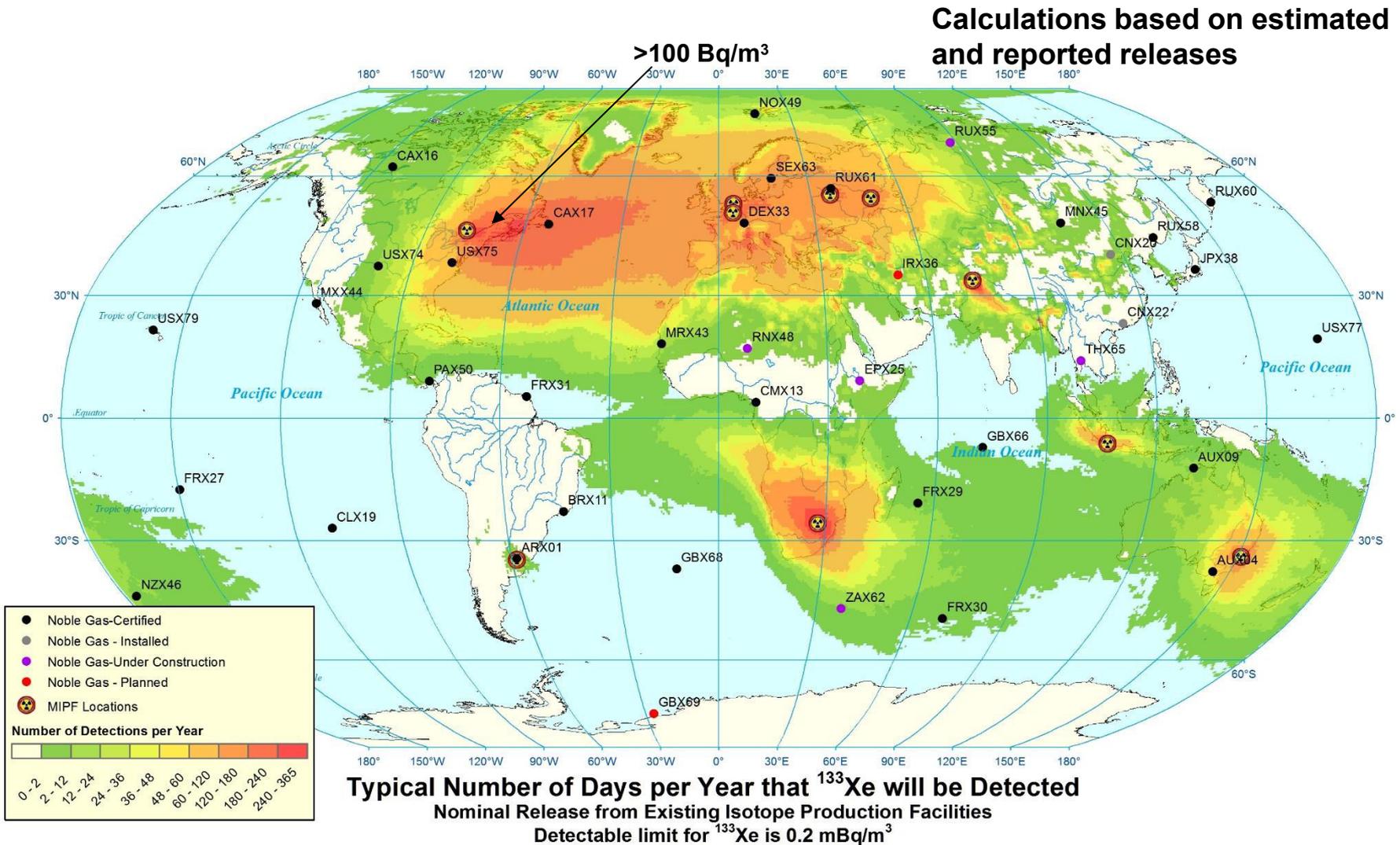
Production of radioactive gases from a nuclear explosion



Man-made backgrounds

- ▶ Over the last several years, we have seen a number of cases in which radionuclide background caused by various man-made phenomena have caused problems with the understanding of radionuclide measurements in the IMS
 - Specifically ^{133}Xe detections have been frequent
- ▶ The situation is getting more frequent and we are in search of tools to try to improve the situation
- ▶ See presentations by Friese, et al.

Map of Xe backgrounds from ^{99}Mo production



Argon-37 for OSI

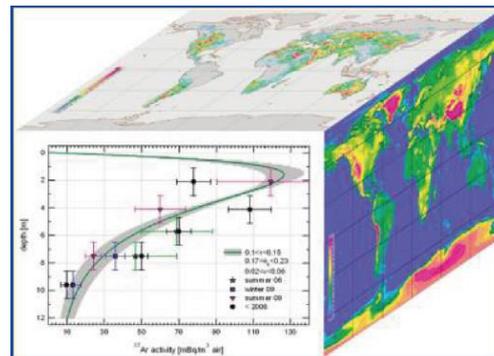
- ▶ Argon-37 has been identified as a nuclear explosion signature for on-site inspection and was discussed at SnT2015 from the China Academy of Engineering Physics.
 - “...a definitive and unambiguous indicator of an underground nuclear explosion.”

- ^{37}Ar signature was verified during the last Chinese UNEs. Subsoil gas sample in two months after the explosion showed that the activity was in the order of $10^2\text{-}10^7\text{Bq/m}^3$.



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- Work by Riedmann and Purschert have also explored the signature, including a study of global backgrounds



Riedmann and Purschert, Env. Sci. Tech. 2011.

A signature to consider for the IMS

- ▶ We should consider augmenting current signatures and bring something else to the table
 - Xe-133 is a good signature; I am not suggesting replacing it
- ▶ A good signature should have a reasonable chance to be detected in the IMS for at least a few weeks
- ▶ Would ideally be based on existing or modification of existing capabilities

Argon-37

Noble gas

Half-life: 35 days

Production: Approximately
 10^{13} - 10^{14} Bq/kt (see below)

The basic advantages of ^{37}Ar for the IMS

- ▶ It has the perfect half-life
 - Long enough to be detectable for a long time, but short enough that it has not built up in the atmosphere
- ▶ The background of ^{37}Ar may be low many places on earth and anyway is most likely different than ^{133}Xe , so that where the ^{133}Xe background is high, ^{37}Ar might not be
- ▶ An explosion is expected to release both ^{37}Ar and ^{133}Xe and therefore, can be detected in coincidence at a station
- ▶ The production yield (Bq/kt) is ‘pretty high’, depending on soil conditions
 - $\sim 10^{13}$ Bq/kt (Gasbuggy, 1971)
 - Up to 10^{14} Bq/kt, depending on rock type (Egnatuk, 2015)

Production yields – ^{37}Ar and ^{133}Xe

▶ Xe-133

- Easy to calculate, based on fission yields
- Cumulative yields maximum at about 3 days (10^{16} Bq)

▶ Ar-37

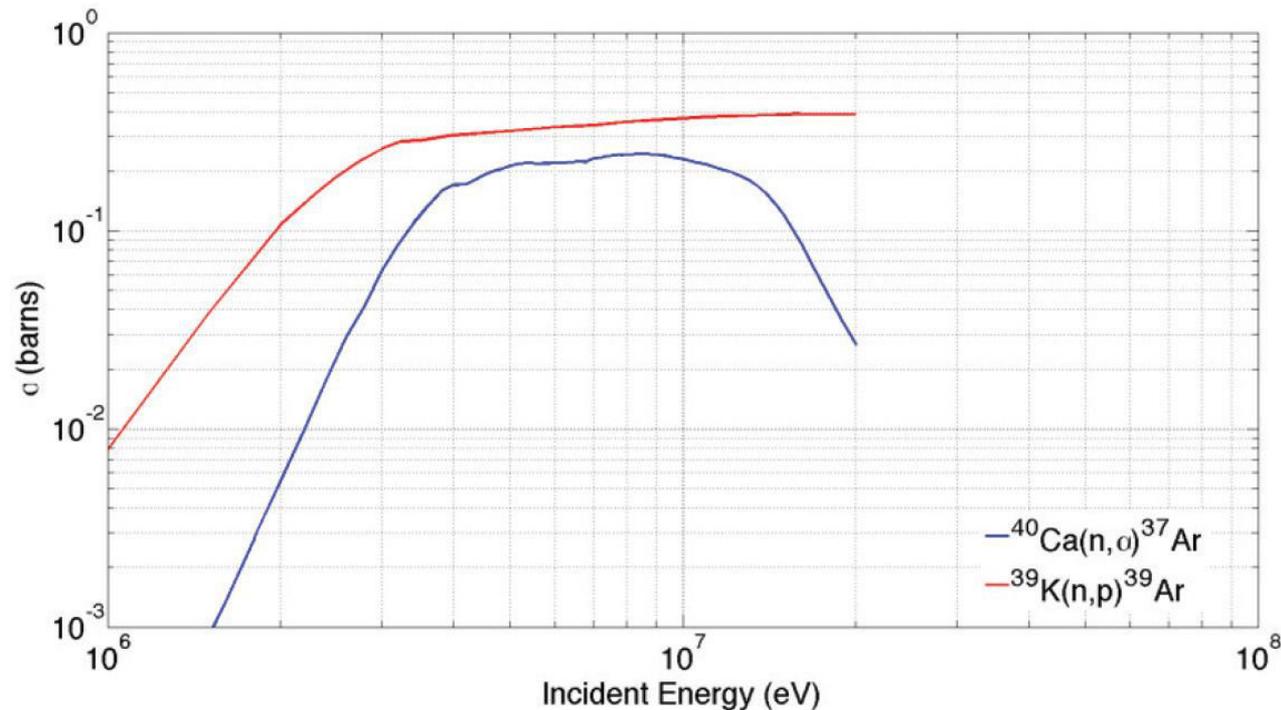
- Estimates range of 10^{13} to 10^{14} Bq/kiloton, depends on
 - The amount of calcium in the soil (assumed a few percent)
 - ◆ By far the largest contributor to the variability
 - The n energy spectrum chosen (Watt spectrum assumed)
 - Details of the simulation (MCNP has been used)

While the cumulative radioxenon and radioargon production are near 10^{14} Bq (Ar-37) to 10^{16} Bq (Xe-133), the amount released can be much less

In principle, the amount of either gas could be “zero”

Source term

- ▶ $10^{13} - 10^{14}$ Bq/kiloton for most soils
 - Created through (n, α) on ^{40}Ca
 - Argon-39 also created via (n,p) on ^{39}K
- ▶ The ^{39}Ar background may be high due to its 269-year half-life
 - See poster by McIntyre

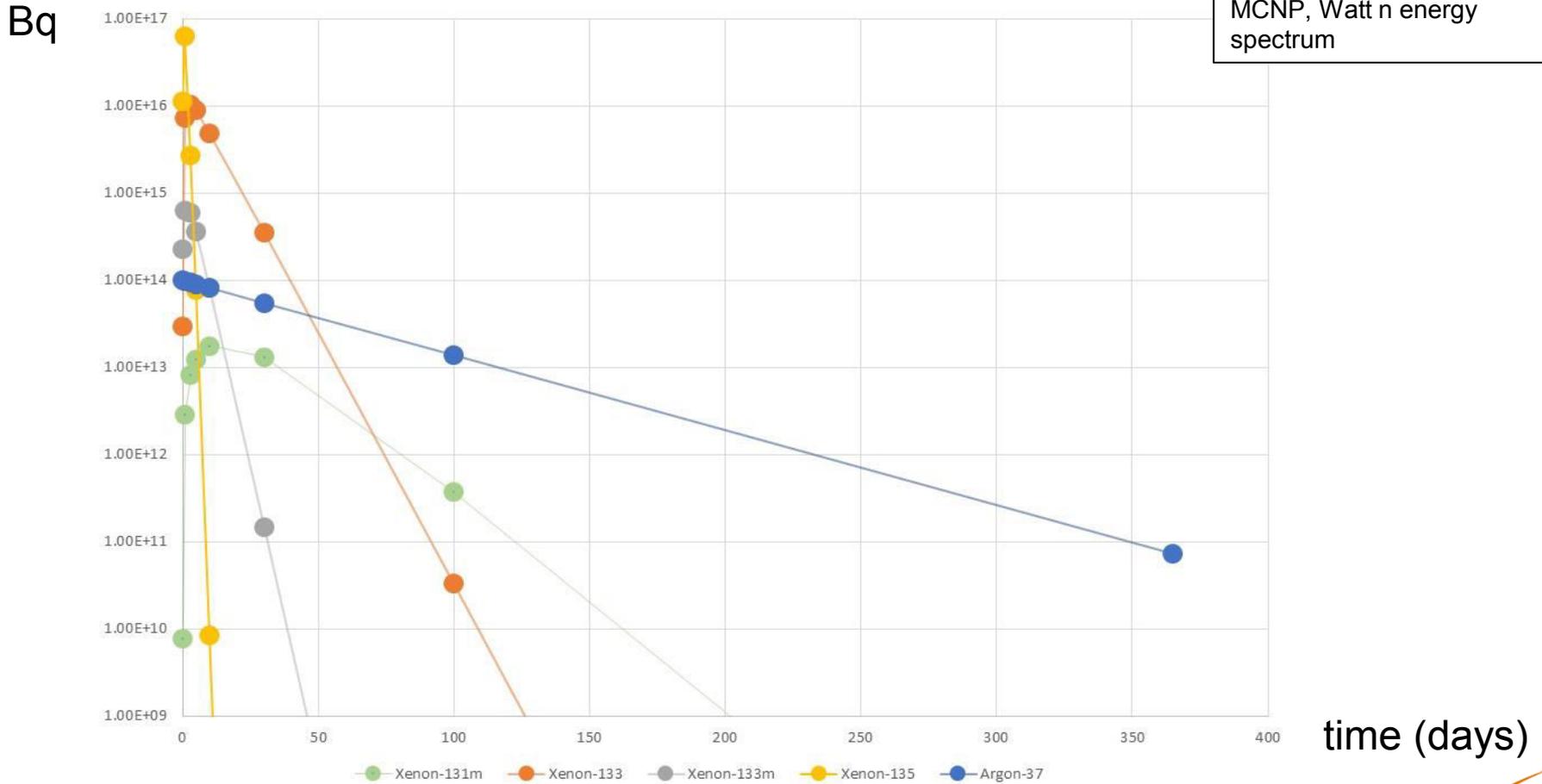


Relative size of the signal from all NGs

^{239}Pu explosion
 10^{14} Bq/kiloton ^{37}Ar , 1% Ca

MCNP, Watt n energy
 spectrum

Radioactive Noble Gases Following a 1 kt blast



Figures-of-merit; comparison to ^{133}Xe

- ▶ Detectability
 - Compare expected signal size to the average background
- ▶ Selectivity
 - Compare expected signal size to background sources and the variability of the background
- ▶ Costs
 - Existing IMS technologies are approximately \$1M+ capital plus \$100K+ annual costs

Figure of Merit (Detectability Ratio)

Expected Concentration / Expected Background

Figures of merit (FOM)

► Define a figure of merit (FOM)

- Expected concentration of an analyte (expected unrelated background) at the measurement point from other processes

Isotope	Half-life (d)	Yield (10 days) Bq	Concentration* (mBq/m3)	Lowest expected** (mBq/m3)	Highest expected*** (mBq/m3)	Average background (mBq/m3)	10 days		
							FOM-	<FOM>	FOM+
Ar-37	35	1.0E+14	1.000	0.1	10	2	10	0.5	0.1
Ar-39	9.8E+04	3.6E+10	3.6E-04	10	10	10	3.6E-05	3.6E-05	3.6E-05
Kr-85	4.0E+03	4.0E+11	4.0E-03	1500	1.5E+06	1.5E+03	2.7E-06	2.7E-06	2.7E-09
Xe-131m	11.9	1.8E+13	0.177	0.1	100	0.2	1.77	0.885	1.8E-03
Xe-133	5.2	4.9E+15	48.500	0.1	1.0E+04	2	485	24.25	4.9E-03
Xe-133m	2.2	8.1E+13	0.813	0.1	100	0.2	8.13	4.065	8.1E-03
Xe-135	0.38	8.5E+09	8.5E-05	0.1	100	0.01	8.5E-04	8.5E-03	8.5E-07

* Assumes 0.1% emission rate

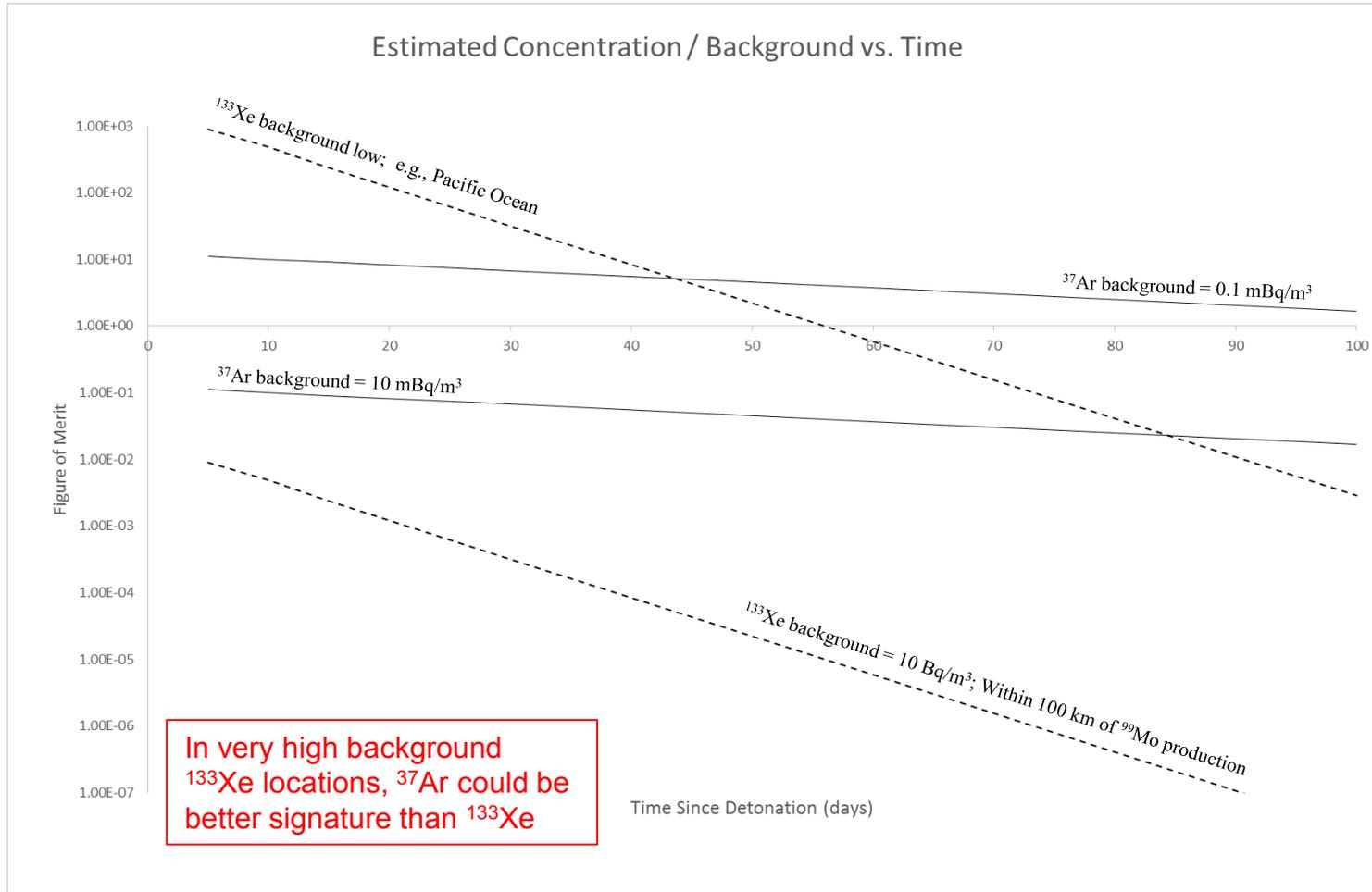
** Lowest concentrations maybe lower, but set at instrumentally-constrained levels (MDCs)

*** Estimate

Low background Average background High background

- Note – some of the inputs to the FOMs are estimates; further work is needed to determine appropriate values

Range of FOMs between Ar and Xe



Interpretation

- ▶ At $t > 80$ days (delayed releases) ^{37}Ar is a better signature than any radioxenon because of its half-life
 - The IMS detected a delayed release from DPRK of ~ 55 days
- ▶ When ^{133}Xe backgrounds are large ($>1 \text{ mBq/m}^3$), ^{37}Ar is probably better than any radioxenon because of its half-life at times >25 days
 - There are some locations, even distant from man-made sources in which the background concentrations of ^{133}Xe exceed 1 mBq/m^3
 - We expect more examples of large background in the future
- ▶ When ^{133}Xe levels are very high ($>10 \text{ mBq/m}^3$), ^{37}Ar is always better than any radioxenon because of its lower background in the atmosphere
 - There are a few locations, close to man-made sources in which the background concentrations of ^{133}Xe exceed 10 mBq/m^3
 - We expect more examples of large background in the future

So, why hasn't this been done?

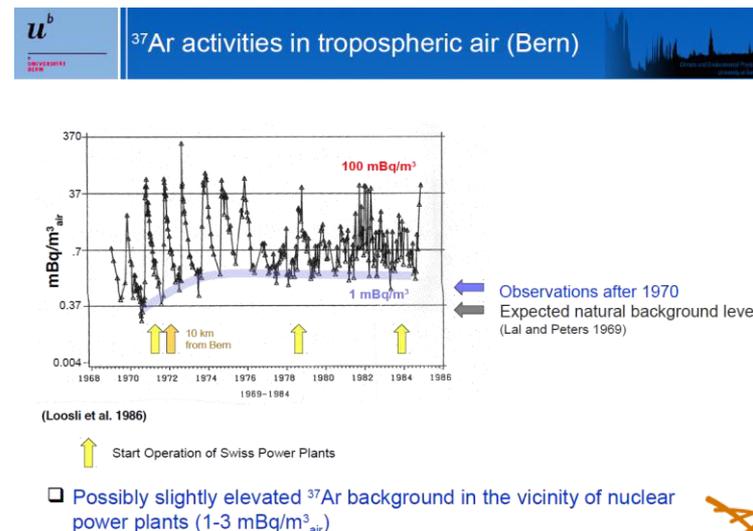
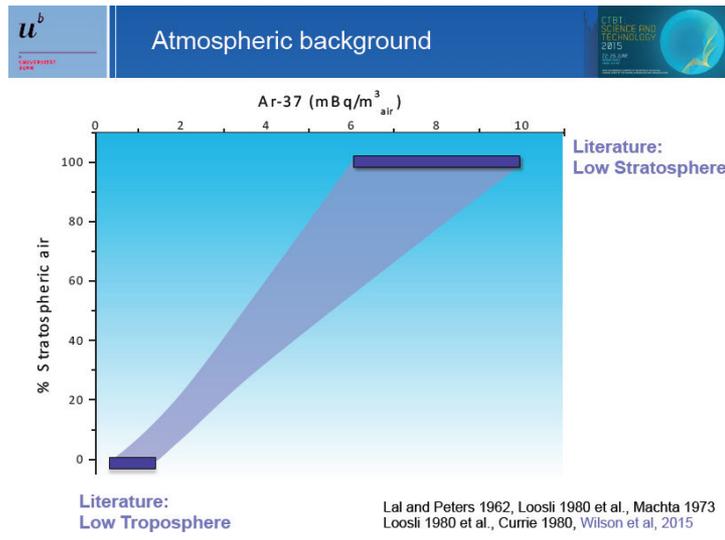
- ▶ Xenon-133 was already being measured and is still an excellent signature
- ▶ Argon-37 is a difficult isotope to measure
 - The only decay mode measurable is a low-energy (2.8 keV) Auger electron
 - This means that a gaseous internal proportional counter is the only likely means for its measurement
 - ◆ A very pure sample is required --- <1 ppm of impurities
 - The gaseous separation between Ar and N₂ & O₂ isn't as simple as heavy noble gases (i.e., Xe)

PRINCIPAL RADIATIONS PRODUCED IN THE DECAY OF ³⁷Ar

Decay Mode	Percent of All Decays	Energy of Auger Electrons (keV)	X-ray (keV)
K	81.5	2.823	0.0
L	8.9	0.270	0.0
K	2.7	0.202	2.621
K	5.5	0.201	2.622
M	0.9	0.018	0.0
K	0.5	0.007	2.816

Information needed

- ▶ Purtschert, et al. have measured ^{37}Ar concentrations at some locations, even some close to reactors and established a relatively low (few mBq/m^3 range) background (but higher in the subsurface)
 - However, a larger data set and map would be beneficial, as well as investigations into release details



Fusion between Xe and Ar data

- ▶ The production of ^{37}Ar and ^{133}Xe come from the same root process (i.e., a nuclear explosion), there are similarities and differences (see presentations by Lowrey) and
 - Releases are similar since they are both noble gases, though some differences in release times might occur
 - The amount of calcium affects the signal size for ^{37}Ar
 - Xe-133 has significant other man-made (and large) sources
- ▶ Hence, if there are correlated detections of both isotopes, the confidence that an event is of a nuclear nature would increase significantly

Conclusions and next steps

- ▶ Argon-37 may be a better signature than Xe-133 in some cases, especially where backgrounds of Xe-133 are high or more than a few months have passed since an explosion

- ▶ Perform a background measurement campaign with defined scientific goals set by the community. Possible areas of concentration include:
 - Measurements at IMS stations
 - A model of nuclear reactors as a source
 - Continued source term models
 - Investigate ^{39}Ar as a background

- ▶ Develop prototype field sensors and concentrate on the following issues (much can be leveraged from OSI systems)
 - Minimization of the use of consumables
 - Increasing sample size to x10 OSI systems
 - Possibly using IMS xenon systems “waste air”
 - Ensuring detector backgrounds are kept low

▶ Backup slides

Argon-37 for the IMS

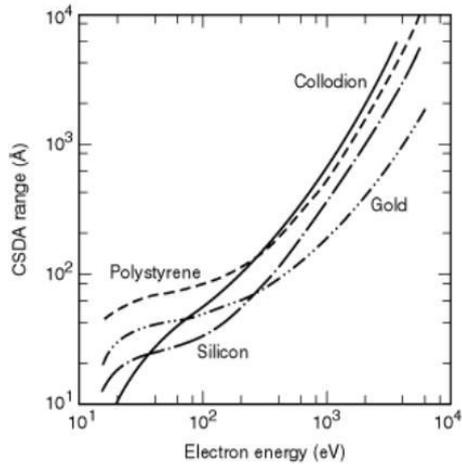
- ▶ While Argon-37 is already a signature for OSI, there are many reasons to examine its use for the IMS
 - The production cross section and half-life makes the isotope perfect
 - Inert, and likely to be emitted from nuclear testing
 - Long half-life (35 d)
 - May help solve background issues associated with man-made processes
 - Correlation to OSI measurements may be needed

Decay of ^{37}Ar

^{37}Ar (35.04 days)

813 keV ν (9.8%)
811 keV ν (90.2%)

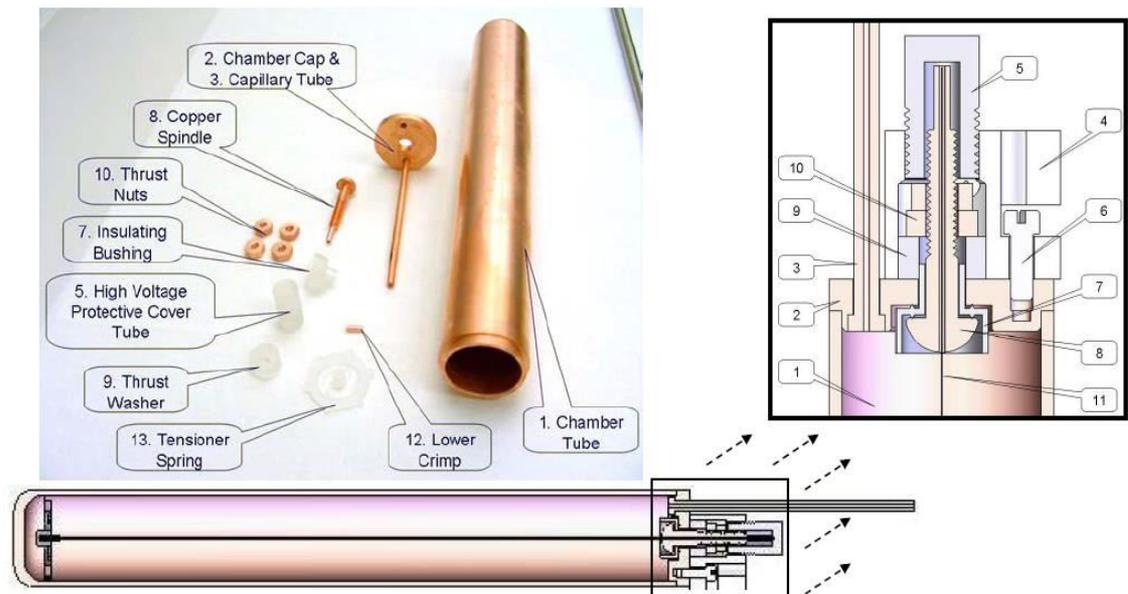
^{37}Cl (stable)



PRINCIPAL RADIATIONS PRODUCED IN THE DECAY OF ^{37}Ar

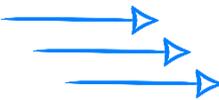
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Attenuation length for
2.8 keV Auger electron
~0.01 cm in Ar

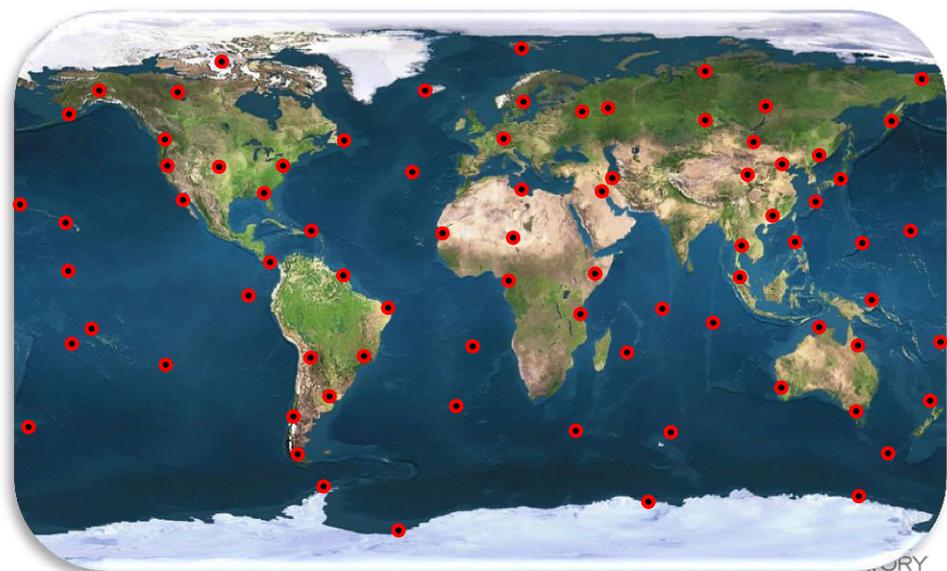
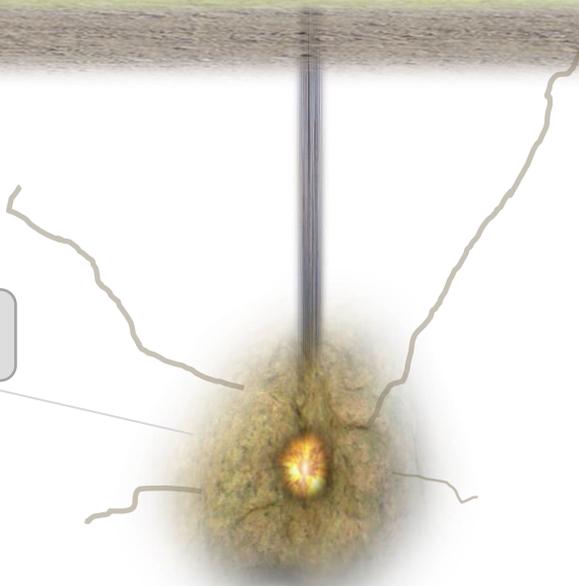


Gas release into the atmosphere and detection in the International Monitoring System (IMS)

Wind



Calcium
Potassium



Detection sensitivity

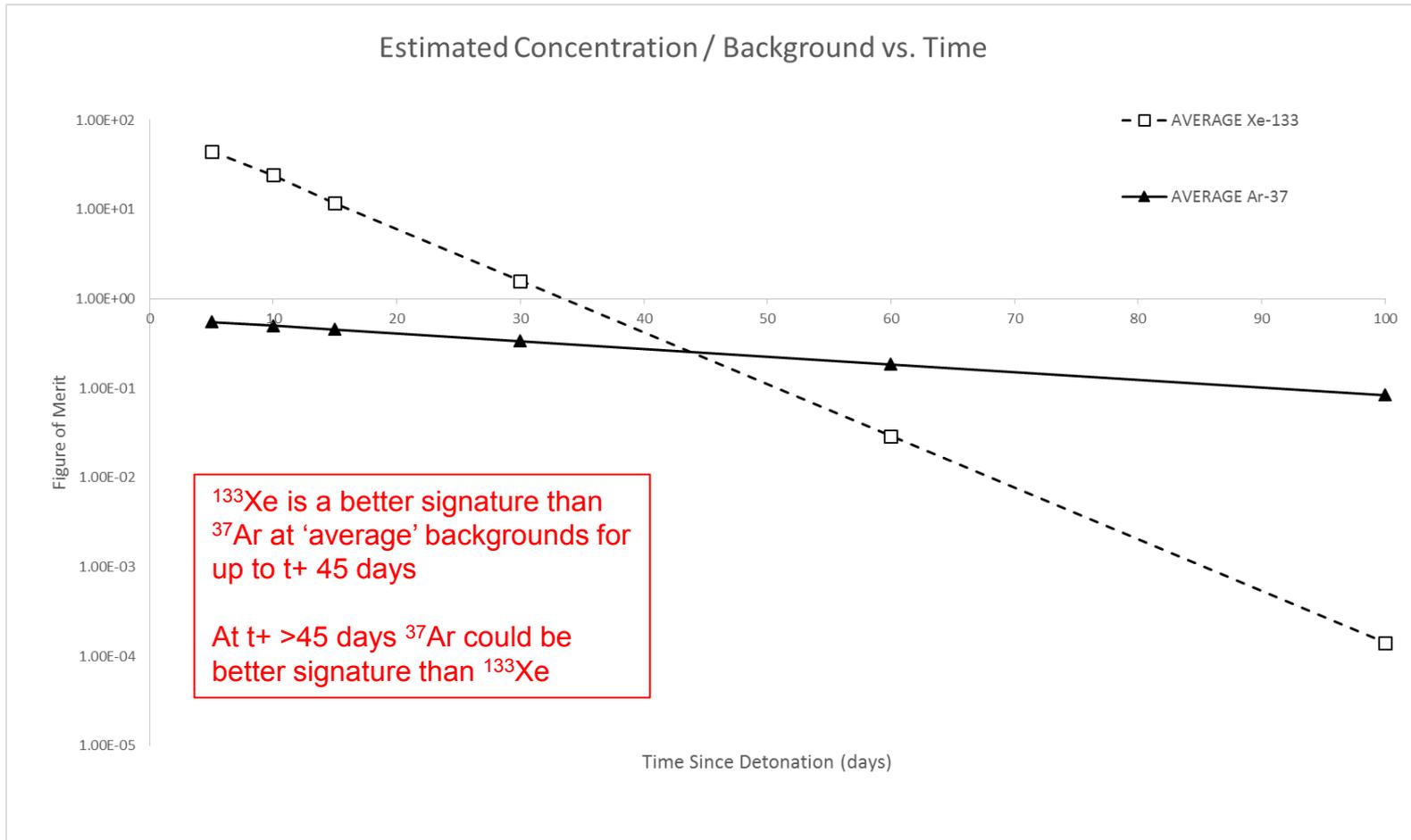
- ▶ Current goals of OSI systems reach MDCs in the range of 10 mBq/m³ (12 hr count)

- Detection levels depend on
 - The count time (t)
 - The amount of sample collected (V)
 - The background (n_b)

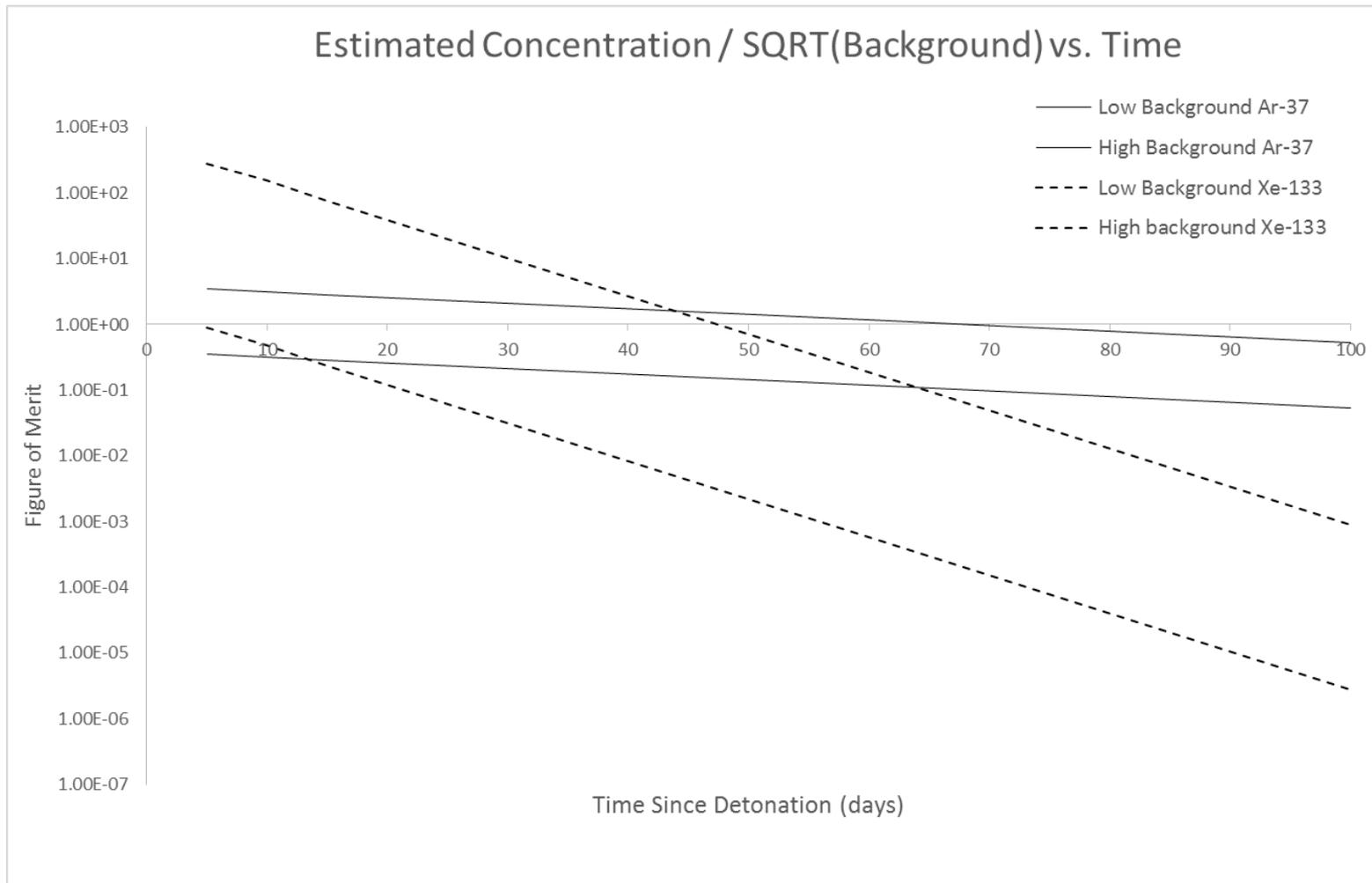
$$\text{MDC} = \frac{4.66 \times \sqrt{n_b \times t}}{V \times \eta \times \varepsilon \times I \times t}$$

- ▶ Lab-based detection sensitivities for *14-day* counts are ~0.05 mBq/m³, for a 1 liter Ar sample (!)
- ▶ For an IMS measurement, increasing the sample size (x4 or more?) and counting time (x4) over the OSI field system
 - These improvements for a 48-hr count could drive the MDC to ~1 mBq/m³ for a field system or better?
 - Exploring the maximum sample size may be the best route

Average case FOM



Alternate FOM (using SQRT of background)



History of ^{37}Ar measurement technology in the last 20 years

- ▶ A significant amount of R&D has been done by the University of Bern since the 1970s
 - The main applications have been for subsurface water tracer studies
 - More recent work by Purschert has been focused on OSI applications
- ▶ China NINT
 - MARDS field system (2008-Present)
- ▶ PNNL
 - ^{37}Ar field system (2015-Present)
 - ^{37}Ar lab capability (2009-Present)