

T1.1-P22

### Did you know?

- **Infrasound propagation:** low frequency acoustic waves (infrasound) are emitted by various sources and can propagate in atmospheric ducts/waveguides, especially in the stratosphere;
- **Weather forecast:** there is a lack of dynamical measurement in the stratosphere to bound the initial conditions of Numerical Weather Prediction (NWP) models.

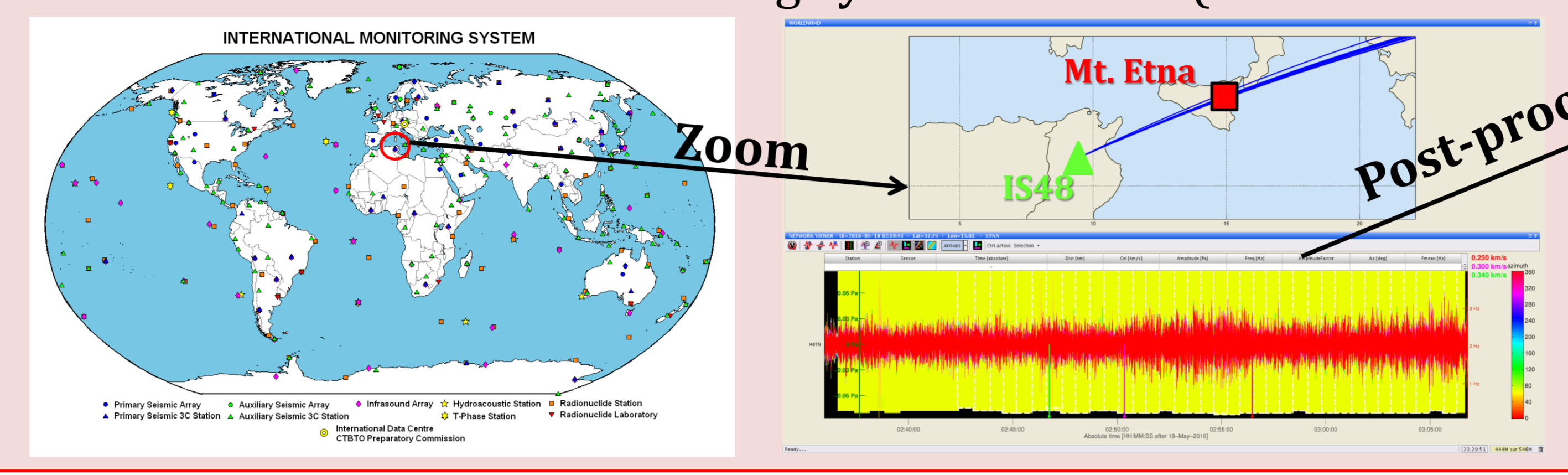
### Our goal:

Use the sensitivity of **infrasound** detections to stratospheric structures for constraining the French **NWP** model ARPEGE<sup>1</sup>.

### Specifying Y

#### Infrasound data available:

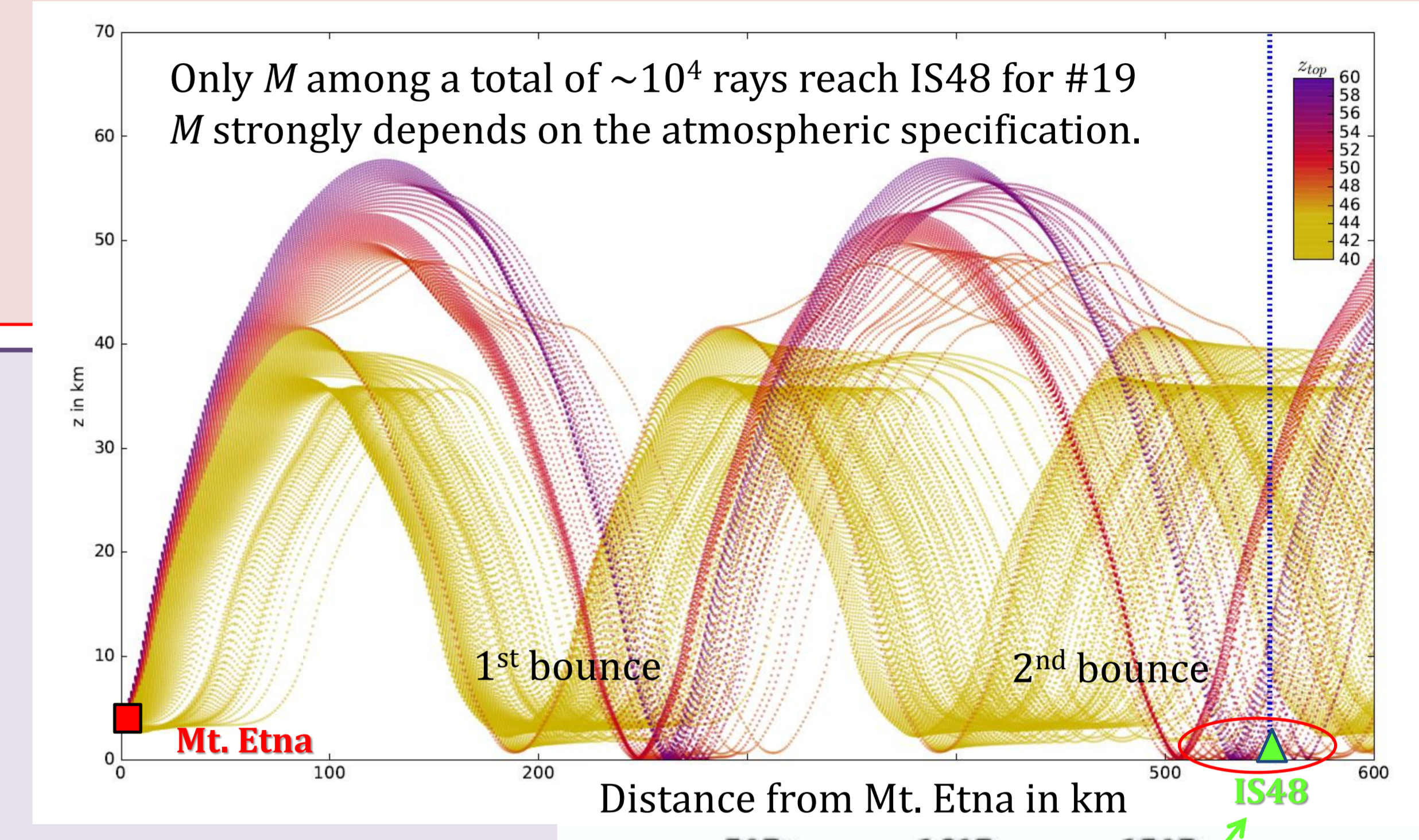
Infrasound detections gathered by the International Monitoring System network (51 infrasound stations IS).



#### Key components of a detection:

- i=1. travel time,
- i=2. back-azimuth angle
- j= 3. trace velocity, ...

Can a given detection be related to a source?



### Specifying X

#### From ARPEGE ensemble analysis (AEARP)

x25 for AEARP

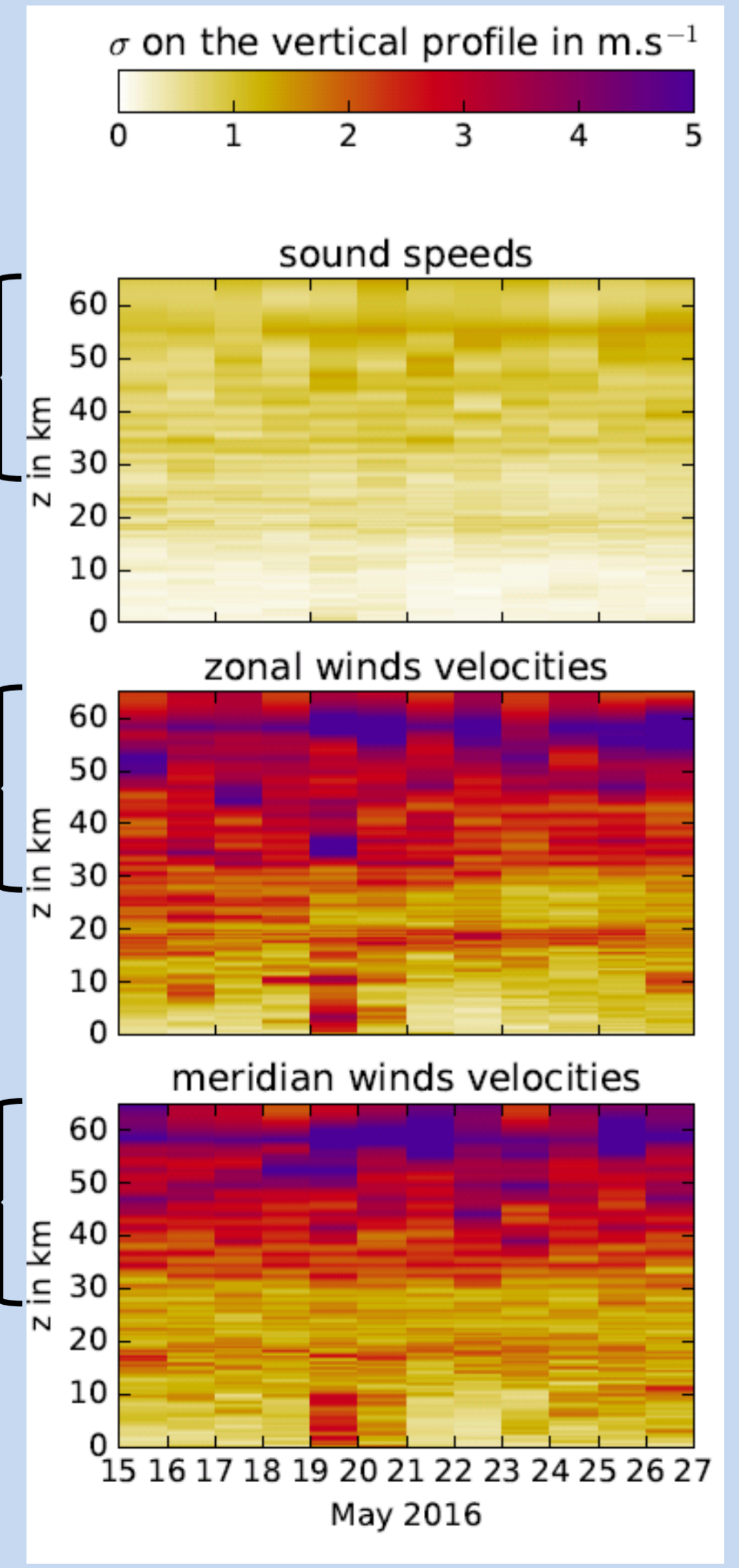
#### to atmospheric specifications (X)

- Extraction of atmospheric quantities: wind components, temperature, pressure and specific humidity from the AEARP grid (reduced Gaussian grid - 42km and 105 hybrid-pressure vertical levels).
- Conversion of these quantities into 3D atmospheric specifications (X) and interpolation onto regularly distributed vertical levels.

Variability in the stratosphere:

- 2-3 m.s<sup>-1</sup> for c(z)
- 10-15 m.s<sup>-1</sup> for U(z)
- 10-15 m.s<sup>-1</sup> for V(z)

The ensemble carries out different scenarios in the stratosphere. Then this variability should be seen on the propagation, especially on the back-azimuth deviations  $\delta\phi$ .



### Selection of 'best' X|Y

$\hat{Y}$  are independent and identically distributed

$$P(X|\hat{Y}) \propto P(Y(X) = \hat{Y}|X)P(X)$$

Those  $X^k$ !

Uniform

$$\exp\left(-\min_{j \in [1;M]} \sum_i \frac{(\hat{y}_i - y_{i,j,k})}{2r_i}\right)$$

No error on Y(X)

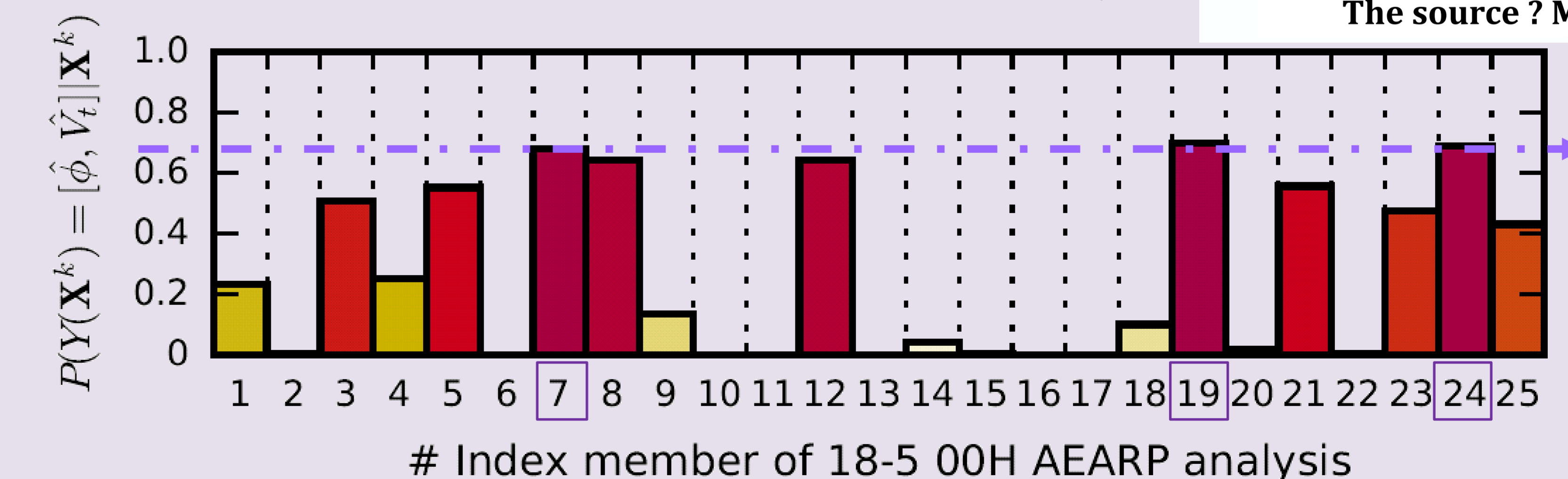
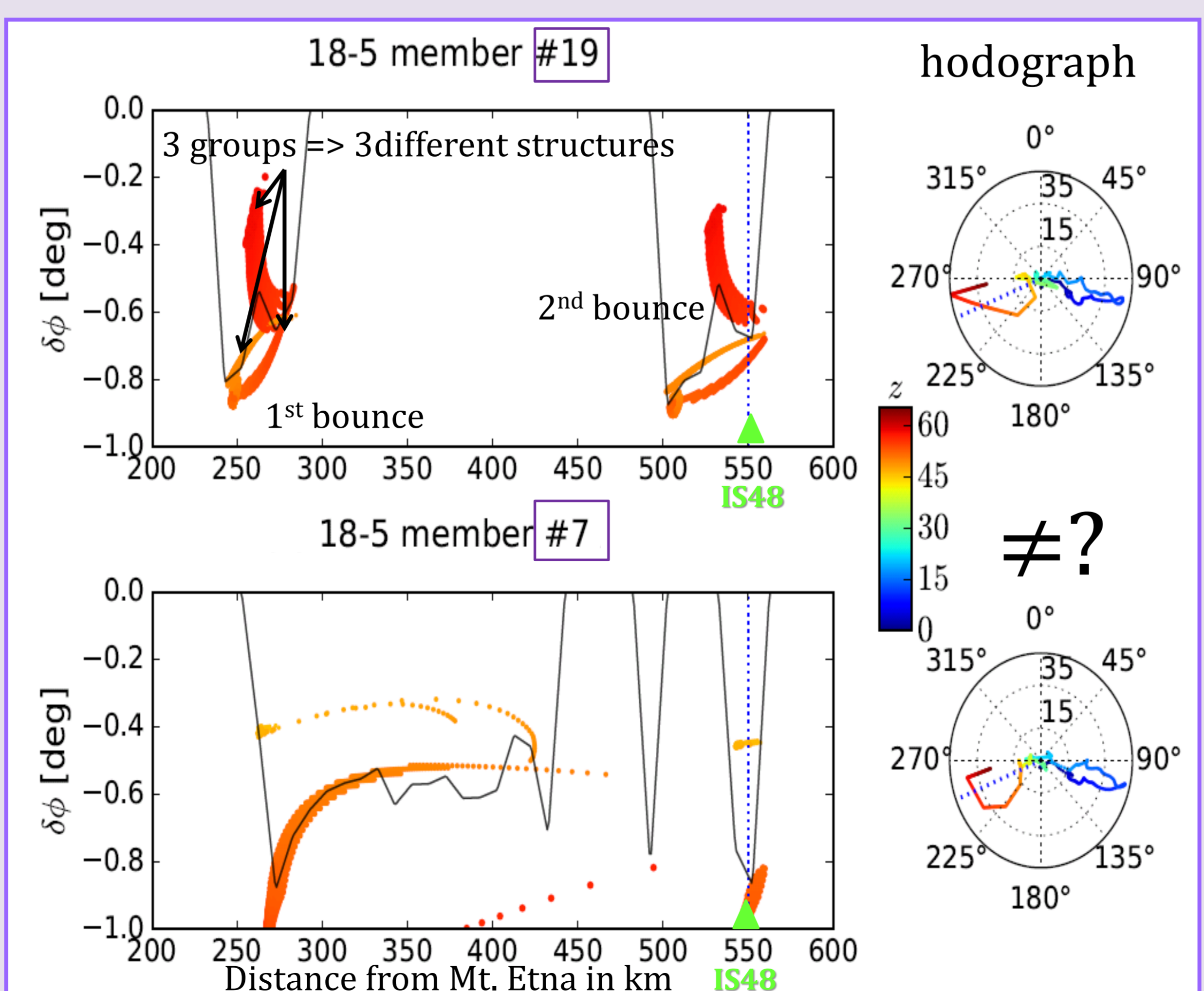
$Y(X^k)$ : which  $X^k$ ?

**Covariance matrix**  
Same for all the members and observations.  
10 s (i=1), 0.5° (i=2) and 5 m.s<sup>-1</sup> (i=3) for IS48.

where  $i, j, k$  denote the components of the observation vector  $Y$ , the ray index and the member.

### Mt. Etna activity on 18<sup>th</sup> May 2016 [02h00 ; 03h00 UTC]

The averaged probability distribution is computed through considering only the backazimuth  $\phi$  and the trace velocity  $V_t$ .



Members #7, #19, #24 lead to quite close effective sound speed profiles but rather different T,U,V, profiles.

### Conclusions:

- While the Bayesian approach provides a ranking  $X^{19} > X^4 > \dots$  of the ensemble members according to some infrasound measurements, there is no physically meaningful methods for decoupling the role of U, V and T in the inversion process;
- However, the most likely members are associated with similar effective sound speed profiles, thereby leading to a possibility of designing an observation operator using the infrasound detections;
- Need to pursue this approach for measuring the impact in assimilation processes.

<sup>1</sup> ARPEGE is the operational global NWP model of French meteorological office. Acknowledgments: to E. Marchetti for the near-field detections and to the European Commission's project ARISE2 (Grant Agreement No. 653980)